

1 a

WE KNOW $l = r^2 \sin^2 \theta \frac{d\phi}{dt}$

$$u^\phi = \frac{d\phi}{dt} = \frac{d\phi}{dt} \frac{dt}{d\tau} = \Omega u^t$$

WE ALSO KNOW $l = \sqrt{12} M$ AT THE ISCO

$$\sqrt{12} M = (6M)^2 \sin^2 \theta \Omega u^t$$

1 FOR $\theta = \pi/2$

$$\Omega = \frac{\sqrt{12}}{36 M u^t}$$

WHAT IS u^t ?

$$g_{\alpha\beta} u^\alpha u^\beta = -1 = g_{tt} u^t u^t + g_{\phi\phi} u^\phi u^\phi$$

$$-1 = -\left(1 - \frac{2M}{6M}\right) (u^t)^2 + (6M)^2 \Omega^2 (u^t)^2$$

$$= -\frac{2}{3} (u^t)^2 + \frac{(6M)^2 12}{6^4 M^2 (u^t)^2} (u^t)^2$$

$$-1 = -\frac{2}{3} (u^t)^2 + \frac{12}{36}$$

$$-1 = -\frac{2}{3} (u^t)^2 + \frac{1}{3}$$

$$-\frac{4}{3} = -\frac{2}{3} (u^t)^2$$

$$u^t = \sqrt{2}$$

← PART OF (b)

$$\Omega = \frac{\sqrt{6}}{36M}$$

$$\left(\Omega^2 = \frac{1}{6^3 M^2} = \frac{M^2}{r^3} \right)$$

AS IN 7.46

b

SOLVED ABOVE

$$u^\alpha = \begin{bmatrix} \sqrt{2} \\ 0 \\ 0 \\ \frac{\sqrt{2}}{36M} \end{bmatrix}$$

$$= \begin{bmatrix} \sqrt{2} \\ 0 \\ 0 \\ \frac{1}{3\sqrt{2}M} \end{bmatrix} = \begin{bmatrix} \sqrt{2} \\ 0 \\ 0 \\ \frac{1}{6\sqrt{2}M} \end{bmatrix} = \begin{bmatrix} \sqrt{2} \\ 0 \\ 0 \\ \frac{\sqrt{2}}{18M} \end{bmatrix}$$



(c)

$$\Omega = \frac{d\phi}{dt} = \frac{\sqrt{61}}{36M}$$

$$d\phi = \frac{\sqrt{61}}{36M} dt$$

$$2\pi = \frac{\sqrt{61}}{36M} t$$

$$t = \frac{36M(2\pi)}{\sqrt{61}}$$

$$= 12\sqrt{61}\pi M = 6^{3/2}(2\pi)M$$

FOR $M = M_0$

$t = 0.0005 \text{ sec} \Rightarrow$ SPEEDY-ASS ORBIT!

UNITS

$$2.998 \times 10^8 \text{ m} = 1 \text{ s}$$

$$6.673 \times 10^{-11} \frac{\text{m}^3}{\text{s}^2} = 1 \text{ kg}$$

SO

$$2.476 \times 10^{-36} \text{ s} = 1 \text{ kg}$$

$$\text{SUN: } M_0 = 2.00 \times 10^{30} \text{ kg}$$

$$M_0 = 5.0 \times 10^{-6} \text{ s}$$

(d)

WE HAVE $\frac{dt}{dr} = \sqrt{2}$ (THAT'S u^t)

$$\text{SO } dr = \frac{dt}{\sqrt{2}}$$

$$r = \frac{36M(2\pi)}{\sqrt{12}}$$

FOR $M = M_0$,

$$r = 0.0003 \text{ sec}$$

YIPERS

$$= 6\sqrt{3}\pi M = 12\sqrt{3}\pi M$$

2

(a) $\nabla_{\omega_1} = -g_{\alpha\beta} p^\alpha u^\beta$

SINCE $p^t = 0$ & $u^r = 0$, ONLY SEE MATERS

$$g_{tt} = -\left(1 - \frac{2M}{6M}\right) = -\frac{2}{3}$$

$$g_{\theta\theta} = (6M)^2$$

$$g_{rr} = \left(1 - \frac{2M}{6M}\right)^{-1} = \frac{3}{2}$$

$$\nabla_{\omega_1} = \frac{2}{3} p^t \sqrt{2}$$

$$p^t = \frac{3\nabla_{\omega_1}}{2\sqrt{2}}$$

NORMALIZE: $g_{\alpha\beta} p^\alpha p^\beta = 0$

RADIAL PHORAM: $p^t = p^r = 0$

$$-\frac{2}{3} \left(\frac{3\nabla_{\omega_1}}{2\sqrt{2}}\right)^2 + \frac{3}{2} (p^r)^2 = 0$$

$$(p^r)^2 = \left(\frac{4}{9}\right) \left(\frac{9}{8}\right) (\nabla_{\omega_1})^2$$

$$p^r = \frac{\nabla_{\omega_1}}{\sqrt{2}}$$

(b) $\sum_{\omega} \cdot P_{\omega} = g_{tt} \xi^t p^t = \left(-\frac{2}{3}\right)(1) \left(\frac{3 \hbar \omega_1}{2\sqrt{2}}\right)$

$$\sum_{\omega} \cdot P_{\omega} = \frac{-\hbar \omega_1}{\sqrt{2}}$$

(c) HERE, $\sum_{\omega} \cdot P_{\omega} = g_{tt} \xi^t p^t = (-1)(1)p^t$

$$-p^t = -\frac{\hbar \omega_1}{\sqrt{2}}$$

$$p^t = \frac{\hbar \omega_1}{\sqrt{2}}$$

SO

$$P_{\omega} = \begin{bmatrix} \hbar \omega_1 / \sqrt{2} \\ \hbar \omega_1 / \sqrt{2} \\ 0 \\ 0 \end{bmatrix}$$

(d) $\omega_2 = \frac{\omega_1}{\sqrt{2}}$

$(-\omega_2 \cdot P_{\omega} = +p^t \text{ SINCE } g_{tt} = -1)$

(e) FROM CLASS: IF Θ , WERE FIXED IN SPACE REL. TO B.H.

$$\frac{\omega_2}{\omega_1} = \sqrt{1 - \frac{2M}{R_1}} = \sqrt{\frac{2}{3}} > \sqrt{\frac{1}{2}}$$

ω_2 IS SMALLER \Rightarrow MORE REDSHIFT WHEN Θ , ORBITS

(f) GREEN $\lambda \approx 5500 \text{ \AA}$

DETECTED $\lambda = \sqrt{2} (5500 \text{ \AA}) = \underline{\underline{7800 \text{ \AA}}}$

λ GOES UP $\sqrt{2}$
IF ω GOES DOWN $\sqrt{2}$

VERY RED,
MAYBE JUST
BEYOND HUMAN
VISION

3 a

$$p^\alpha = \begin{bmatrix} M \\ 0 \\ 0 \\ 0 \end{bmatrix}$$

b

$$\sum_w p_w^\alpha = g_{tt} \dot{\Sigma}^t p^t = \underline{-M}$$

c

$$\sum_w p_w^\alpha = -m = g_{tt} \dot{\Sigma}^t p^t$$

$$-m = -\left(1 - \frac{2M}{r}\right) (1) p^t$$

$$-m = -\frac{2}{3} p^t$$

$$\underline{p^t = \frac{3}{2} m}$$

@ r=6M

$$g_{tt} = -\left(1 - \frac{2M}{6M}\right) = -\frac{2}{3}$$

$$g_{rr} = \left(1 - \frac{2M}{6M}\right)^{-1} = \frac{3}{2}$$

$$p^\alpha = m \begin{bmatrix} 3/2 \\ 1/\sqrt{3} \\ 0 \\ 0 \end{bmatrix}$$

For p^r , normalize:

$$-m^2 = g_{tt} (p^t)^2 + g_{rr} (p^r)^2$$

$$p^r = \sqrt{\frac{-m^2 - g_{tt} (p^t)^2}{g_{rr}}}$$

$$p^r = m \sqrt{\frac{-1 - \left(-\frac{2}{3}\right) \left(\frac{3}{2}\right)^2}{3/2}}$$

$$= m \sqrt{\frac{\frac{3}{2} - 1}{3/2}}$$

$$= m \sqrt{\left(\frac{1}{2}\right) \left(\frac{2}{3}\right)}$$

$$p^r = \frac{m}{\sqrt{3}}$$

d

$$E = -u_{\mu 1} \cdot p^\mu \quad \text{SINCE } p^\theta = 0 \text{ \& } u_1^r = 0, \\ \text{ONLY } t \text{ MATTERS}$$

$$\text{FROM 1, WE HAVE } u_1^t = \sqrt{2} \quad g_{tt} = -\frac{2}{3}$$

$$E = +\frac{2}{3} (\sqrt{2}) p^t = \frac{2}{3} (\sqrt{2}) \frac{3}{2} m = \boxed{\sqrt{2} m = E}$$

e M ∇_0 THE PARTICLE IS AT REST
WRT σ_1 NOW

$$\textcircled{f} \quad \boxed{\Delta E = (\sqrt{2} - 1) m}$$

$$\textcircled{14} \quad \text{WE HAVE } E_{\text{phot}} = \hbar \omega_1 = (\sqrt{2} - 1) m$$

$$\textcircled{a} \quad \text{FROM } \textcircled{a}, \quad P^{\pm} = \frac{3 \hbar \omega_1}{2 \sqrt{2}} = \frac{3(\sqrt{2}-1)}{2 \sqrt{2}} m = P^{\pm}$$

$$P^{\mp} = \frac{\sqrt{3}}{12M} \hbar \omega_1 = \frac{\sqrt{3}(\sqrt{2}-1)}{12M} m = P^{\mp}$$

$$\textcircled{b} \quad \sum_{\omega} \cdot P_{\omega \text{ phot}} = \text{const} = 3 \pm \frac{3}{3} P_{\text{phot}}^{\pm}$$

$$= \left(-\frac{2}{3} \right) (1) \left(\frac{3(\sqrt{2}-1)}{2 \sqrt{2}} m \right) = -\frac{\sqrt{2}-1}{\sqrt{2}} m$$

$$\textcircled{c} \quad \infty, \quad \sum_{\omega} \cdot P_{\omega \text{ phot}} = (-1)(1) P_{\text{phot}}^{\pm}$$

$$P_{\text{phot}}^{\pm} = +\frac{\sqrt{2}-1}{\sqrt{2}} m \quad E_{\text{phot}} = -\frac{4}{3} P_{\text{phot}}^{\pm}$$

$$\text{SO } \boxed{E_{\text{phot}} = \frac{\sqrt{2}-1}{\sqrt{2}} m \text{ AT } \infty}$$

$$= 0.3 m$$

WAY BETTER THAN
FUSION

$$\textcircled{c} \quad kT \sim \Delta E = (1 + \sqrt{2}) m$$

$$T \sim \frac{(1 + \sqrt{2}) m}{k}$$

$$\textcircled{d} \quad \text{FOR } m = 1.67 \times 10^{-27} \text{ kg}$$

$$T \sim \frac{(1 + \sqrt{2})(1.67 \times 10^{-27} \text{ kg})(3 \times 10^8 \frac{\text{m}}{\text{s}})^2}{1.38 \times 10^{-23} \text{ J/K}}$$

$$T \sim 3 \times 10^{12} \text{ K} \quad \text{THAT'S HOT!}$$

PEAK \rightarrow IS GAMMA RAY

e) CONT'D

THAT'S REALLY HOT

WE'VE OVER SIMPLIFIED, OF COURSE ← AND THIS IS
JUST THE INNER
EDGE OF THE
DISK...

T NP-100'S OF MILLIONS ARE

OBSERVED SOMETIMES,

AND SOME MORE DETAILED
THEORY SUGGESTS INNER
DISK T OF 10" ISH K

f)

 $\sqrt{2}$

LOWER ⇒ STILL A BADASS GAMMA RAY!